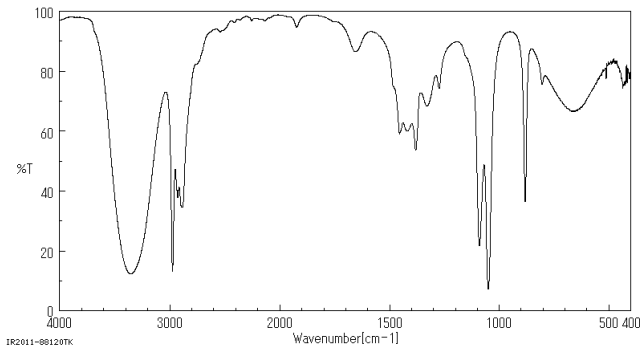


# Chemistry 2000 Slide Set 6: Vibrational spectroscopy of polyatomic molecules

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January 14, 2020

# Example: IR spectrum of liquid ethanol





Source: Spectral Database of Organic Compounds, [http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre\\_index.cgi](http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre_index.cgi), Jan. 16, 2013

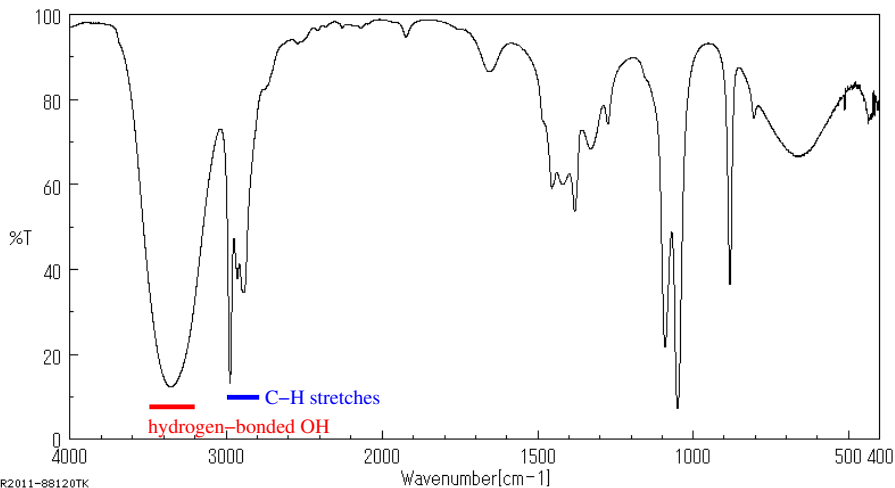
Note: The wavenumber axis often runs backward, as shown here.

# Infrared spectroscopy and the identification of compounds

- One important application of spectroscopy (in general) is for the identification of unknown compounds.
- Certain bonds in organic molecules are associated with characteristic IR bands in specific spectral regions:

Bond	Spectral region/ $\text{cm}^{-1}$
	2800–3000
 (including aromatic CH)	3000–3200
O–H (non-hydrogen-bonded)	3500–3700 (sharp)
O–H (hydrogen-bonded)	3200–3500 (broad)

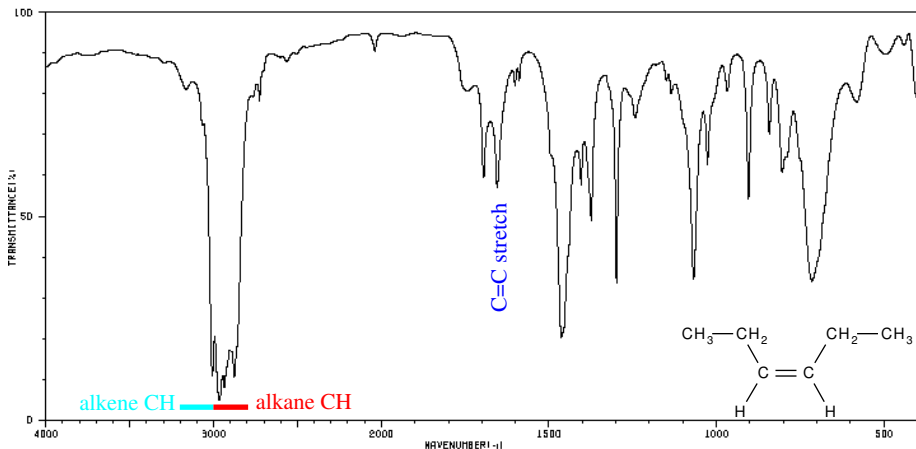
## Example: The IR spectrum of ethanol



# Alkene and alkyne carbon-carbon bond stretches

<b>Bond</b>	<b>Spectral region/cm<sup>-1</sup></b>
C=C	1640–1675 (sometimes)
C≡C	1950–2300 (sometimes)

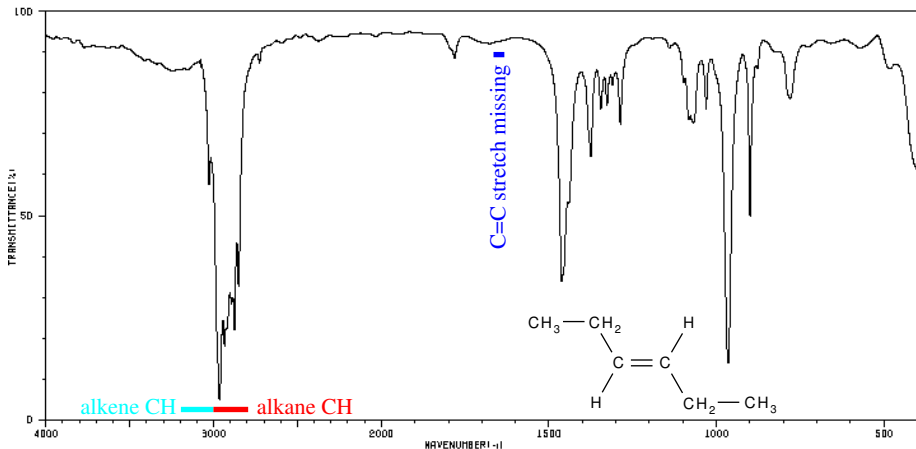
# Example: IR spectrum of liquid *cis*-3-hexene



Spectrum source: Spectral Database of Organic Compounds,

[http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre\\_index.cgi](http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre_index.cgi), Jan. 20, 2013

# Example: IR spectrum of liquid *trans*-3-hexene



Spectrum source: Spectral Database of Organic Compounds,

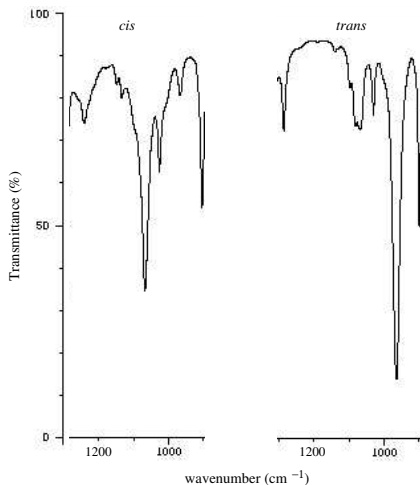
[http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre\\_index.cgi](http://riodb01.ibase.aist.go.jp/sdbs/cgi-bin/cre_index.cgi), Jan. 20, 2013

# The fingerprint region of the spectrum

- The region from 900 to 1300  $\text{cm}^{-1}$  is called the **fingerprint region** of the IR spectrum.
- In this region, we typically find many peaks arising from various low-energy stretching and bending motions of the molecules.
- Very difficult to assign peaks in this region **but** they are very different even for closely related compounds
- Used for confirmation that a particular (known) compound has been isolated



# Example: Fingerprint regions of *cis*- and *trans*-3-hexene compared



# Review: Molecular dipole moments

- A **bond dipole** is a slight separation of charge between two non-identical atoms connected by a bond.
- The size of the bond dipole is proportional to the amount of charge separation and to the bond length.
- The **dipole moment** of a molecule is the **vector sum** of the bond dipoles.
- A polar molecule has a non-zero dipole moment.
- Examples:  $\text{CO}_2$ ,  $\text{H}_2\text{O}$

# Normal modes

- Except in diatomics, molecular vibrations generally involve motions of several atoms, i.e. more than one bond is deformed at a time.
- The vibrational modes must conserve overall molecular momentum.
- We can choose vibrational modes that are independent motions, called **normal modes**.

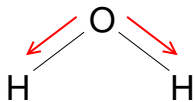
# Number of normal modes

- A molecule made up of  $N$  atoms can move in  $3N$  different ways (one direction of motion per atom per Cartesian axis).
- 3 of these motions are associated with the translational motion of the molecule as a whole.
- A **nonlinear** molecule has 3 modes associated with rotation of the molecule as a whole.
- The remaining  $3N - 6$  modes of a nonlinear molecule are the normal modes of vibration.
- A **linear** molecule only has 2 rotational modes.
- The remaining  $3N - 5$  modes of a linear molecule are the vibrational normal modes.

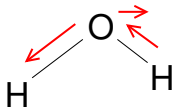
# Normal modes of H<sub>2</sub>O

- $N = 3$  atoms, nonlinear molecule

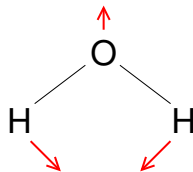
⇒ 3 normal modes



Symmetric stretch



Asymmetric stretch



Bend

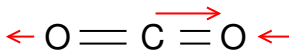
# Normal modes of CO<sub>2</sub>

- $N = 3$  atoms, linear molecule

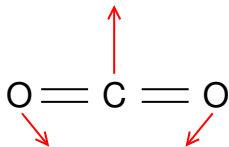
⇒ 4 normal modes



Symmetric stretch



Asymmetric stretch



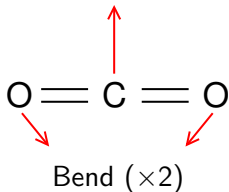
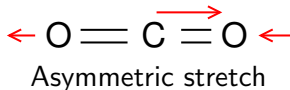
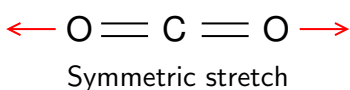
Bend ( $\times 2$ )

# Selection rule

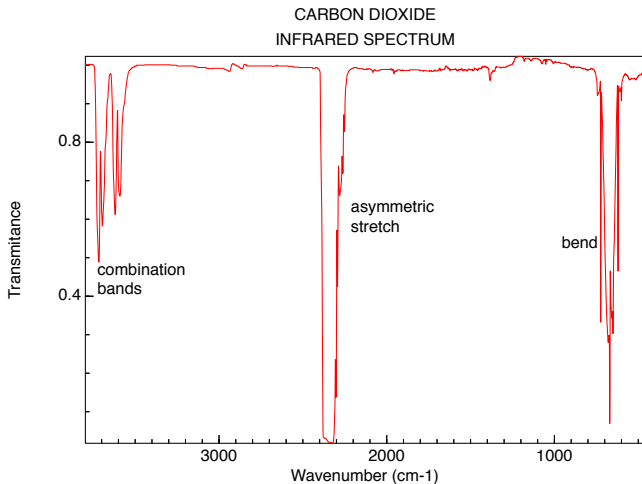
- A **selection rule** is a rule that tells us when a particular kind of spectroscopic event can occur.
- In IR absorption spectroscopy, the key selection rule is that the **dipole moment of the molecule has to change during the vibration**.
- A normal mode that can absorb an IR photon is said to be **IR active**.

# Normal modes of CO<sub>2</sub> in IR spectroscopy

Which of these modes are IR active?



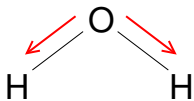


IR spectrum of CO<sub>2</sub>

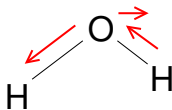
NIST Chemistry WebBook (<https://webbook.nist.gov/chemistry>)

Normal modes of H<sub>2</sub>O in IR spectroscopy

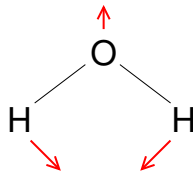
Which of these modes are IR active?



Symmetric stretch



Asymmetric stretch



Bend

## Application: Earth's heat balance

- Energy from the Sun mostly arrives at the Earth in the form of visible light.

Note that the atmosphere is essentially transparent at optical wavelengths.

- The Earth reflects some of that energy (esp. snow and ice at poles), but absorbs a lot of it.

Averaged over the whole planet, about 30% of the light coming in is just reflected back to space.

- The planet radiates mostly in the infrared (blackbody radiation).
- The atmosphere contains many gases that absorb in the infrared, so some of the radiation from the Earth is absorbed in the atmosphere, but then **what happens to the energy captured by the atmosphere?**

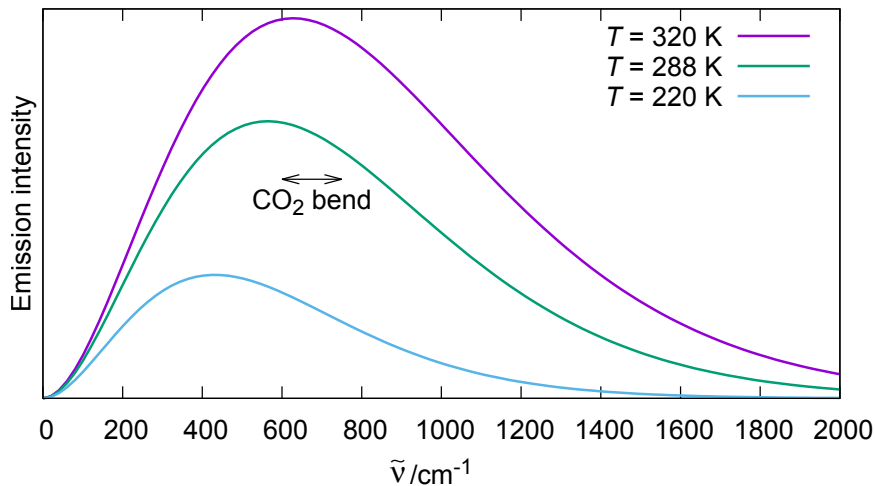
# Application: Earth's heat balance

## Greenhouse gases

- When a gaseous molecule becomes vibrationally excited by absorbing infrared radiation, the excess vibrational energy can be converted to translational kinetic energy during collisions.
- Energy is constantly redistributed in collisions and other energy-transfer processes.
- A gas at temperature  $T$  also emits “blackbody” radiation.

## Application: Earth's heat balance

## Blackbody curves



# Application: Earth's heat balance

## Greenhouse gases

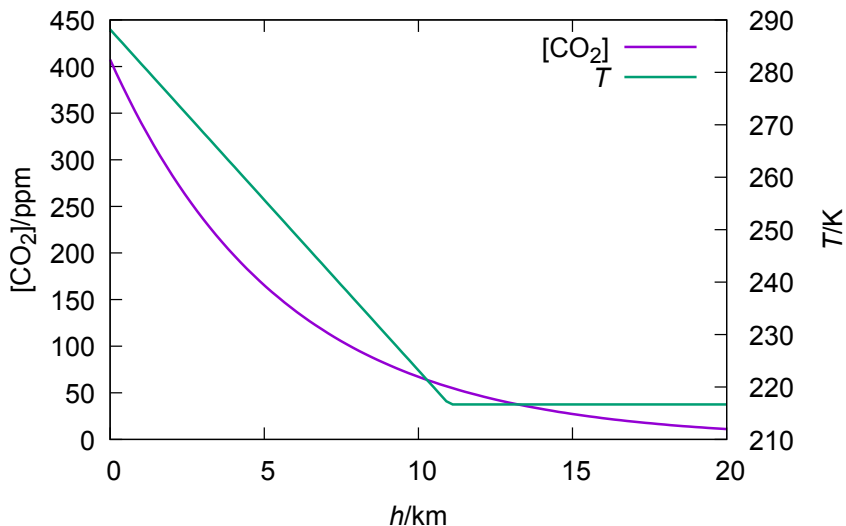
- $\text{N}_2$ ,  $\text{O}_2$  and Ar, the major components of the atmosphere, don't absorb in the IR. (Why?)
- The next two most common components of the atmosphere, water and carbon dioxide do absorb in the IR.
- Gases that absorb in the IR are called **greenhouse gases**.
- The atmospheric water content is set by the balance of evaporation and precipitation, which depends on the atmospheric temperature. It is a **responding variable**.
- We worry a lot about  $\text{CO}_2$  because we are adding a lot of it to the atmosphere, which affects energy transfer through the atmosphere.

# Application: Earth's heat balance

## Photons reabsorbed vs lost to space

- At lower altitudes, photons emitted at wavelengths that  $\text{CO}_2$  can absorb travel only a short distance (a few meters) before they are in fact absorbed by a  $\text{CO}_2$  molecule.  
Similar statements could be made about other greenhouse gases in their respective absorption ranges.
- Absorption of IR photons slows the migration of heat through the atmosphere.
- Near the top of the atmosphere, where the pressure of  $\text{CO}_2$  is low, there is a much larger probability that a photon emitted toward space will actually escape without being reabsorbed.
- Important fact: At those altitudes, the atmosphere is a lot cooler.

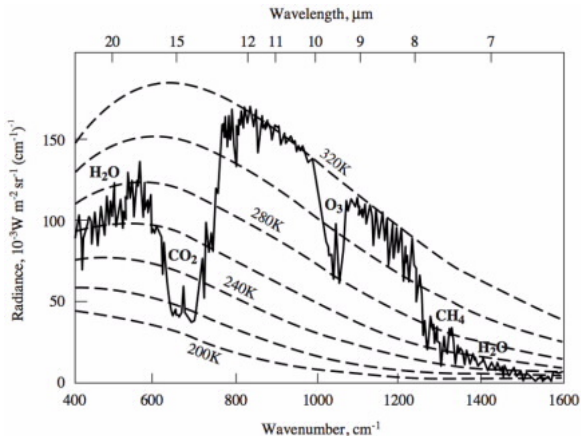
## Application: Earth's heat balance

CO<sub>2</sub> concentration and temperature vs altitude



# Application: Earth's heat balance

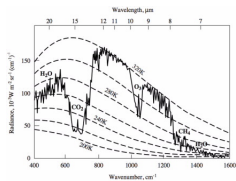
Earth emission spectrum (taken over North Africa)



Hanel et al., *J. Geophys. Res.* **77**, 2829 (1972)

# Application: Earth's heat balance

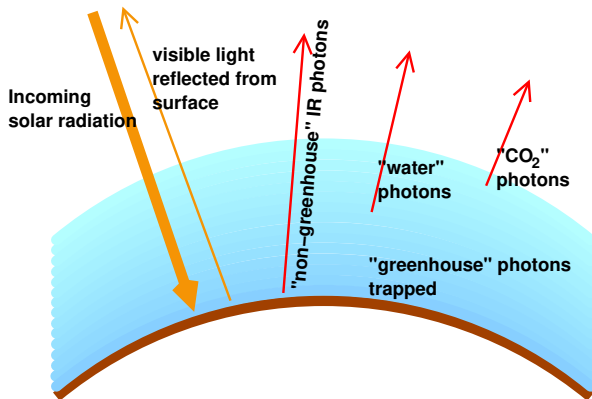
## Interpretation of Earth's emission spectrum



- Emission from CO<sub>2</sub> comes from high in the atmosphere (where it's cool and there isn't much CO<sub>2</sub> to block the outgoing IR photons).
- Emission from water comes from lower down (where it's not quite as cool, since condensation of water prevents it from getting too high in the atmosphere).

# Application: Earth's heat balance

## Interpretation of Earth's emission spectrum



# The wrap-up

- Greenhouse gases like  $\text{CO}_2$  slow the escape of heat from the atmosphere to space.
- A simple analogy is that the atmosphere acts like a blanket.
- This is not inherently a bad thing. The planet would be a lot colder (average surface temperature of about 255 K, or  $-18^\circ\text{C}$ ) if there were no greenhouse effect.
- Adding greenhouse gases to the atmosphere is analogous to making the blanket denser, resulting in a higher temperature under the blanket.

# Application: Earth's heat balance

## Carbon dioxide

- From 1959 to 2019, the CO<sub>2</sub> concentration in the atmosphere measured at the Mauna Loa observatory has risen from an annual average value of 316 ppm to 411 ppm, an **increase of 30%**.
- The rate of increase in the CO<sub>2</sub> concentration is also rising, from about 0.6 ppm y<sup>-1</sup> in the early 1960s to about 2.6 ppm y<sup>-1</sup> now.
- Warming induced by greenhouse gas emissions is a self-reinforcing problem:
  - It increases the amount of water vapor in the atmosphere.
  - On average, less of the planet is covered with ice.
  - Melting permafrost releases methane, a very powerful greenhouse gas.
  - ...
- **There is no escaping the physics: adding greenhouse gases to the atmosphere heats up the planet.**